

3CSL

From XDSwiki

HasA/R (PDB id 3CSL (<http://www.pdb.org/pdb/explore/explore.do?structureId=3CSL>)) is a complex of a 22-stranded beta-barrel outer membrane protein (HAsR, 865 residues), its hemophore (HasA, 206 residues), and heme. The structure and its biological implications are described in "Heme uptake across the outer membrane as revealed by crystal structures of the receptor-hemophore complex" (Krieg, S., Huché, F., Diederichs, K., Izadi-Pruneyre, N., Lecroisey, A., Wandersman, C., Delepelaire, P., Welte, W. (2009), Proc. Nat. Acad. Sci. Vol. 106 pp. 1045-1050.)

3-wl SeMet-MAD data were collected at beamline X06SA of the SLS in November 2006 on a MarCCD detector. HasA/R crystallizes in spacegroup F222; cell parameters are $a=157\text{\AA}$, $b=163\text{\AA}$, $c=596\text{\AA}$. There are 2 complexes per ASU. Data to about 3.0\AA could be collected from this crystal, but the anomalous data are useful to about 5\AA only. The ordered part of HasR has residues 112-865 and harbours 9 SeMet residues. The ordered part of HasA has 173 residues, one of which is SeMet - but that is mostly disordered.

These MAD data, giving a structure with an average B of 100\AA^2 , constitute a project that is challenging for humans, and currently too difficult for automatic methods of structure solution and model building. The deposited 3CSL structure was not obtained from these MAD data alone, but the model was actually refined against slightly better (2.7\AA) data collected on a native crystal at the ESRF.

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XDS data reduction of high-remote, peak and

inflection

The script `generate_XDS.INP` may be used to get a suitable first XDS.INP file for each of the three wavelengths. Unfortunately the beamline software did not put the correct X and Y position of the direct beam into the header. So you will have to find this yourself, using `adxv` or `XDS-viewer`. Or just use:

```
ORGX= 1536 ORGY= 1520
```

The other thing that you might want to try yourself, or just fill in, is

```
VALUE_RANGE_FOR_TRUSTED_DETECTOR_PIXELS=8000. 30000. ! often 8000 is ok
```

instead of the the default (7000. 30000.). This results in a good mask for the beamstop shadow.

It turns out that the spot shapes are actually so irregular that XDS stops after the `IDXREF` step, with a long warning message. This is because it cannot index (within default error margins) enough reflections (50% is the cutoff). When that occurs, one simply continues with the step after `IDXREF`:

```
JOBS= DEFPIX INTEGRATE CORRECT
```

Other than that, the three MAD wavelengths can be processed once with default parameters, as written into `XDS.INP` by `generate_XDS.INP`. This data reduction therefore proceeds in spacegroup `P1`, but the correct spacegroup (22) is identified by `CORRECT`.

Optimization: after this first data reduction pass, I use the "post-refined" geometric parameters, and the correct spacegroup (as given in `CORRECT.LP`, and written to `GXPARM.XDS`), for a second pass. Thus I need to

```
mv GXPAM.XDS XPARAM.XDS
```

and modify `XDS.INP` to read

```
JOBS= INTEGRATE CORRECT
```

Afterwards, another `xds_par` run gives the final intensities. Repeating this optimization sometimes helps.

Peak

360 frames (0.5° oscillation) at the peak wavelength were collected after the high-remote data. They can be downloaded from here (<ftp://turn5.biologie.uni-konstanz.de/pub/datasets/3csl-pk.tar>) (1.9 Gb). This peak dataset is somewhat difficult to index; if the results are really bad (e.g. distance refining far away from 370 mm) with the default 180 frames, then just try with 90 or 270 frames.

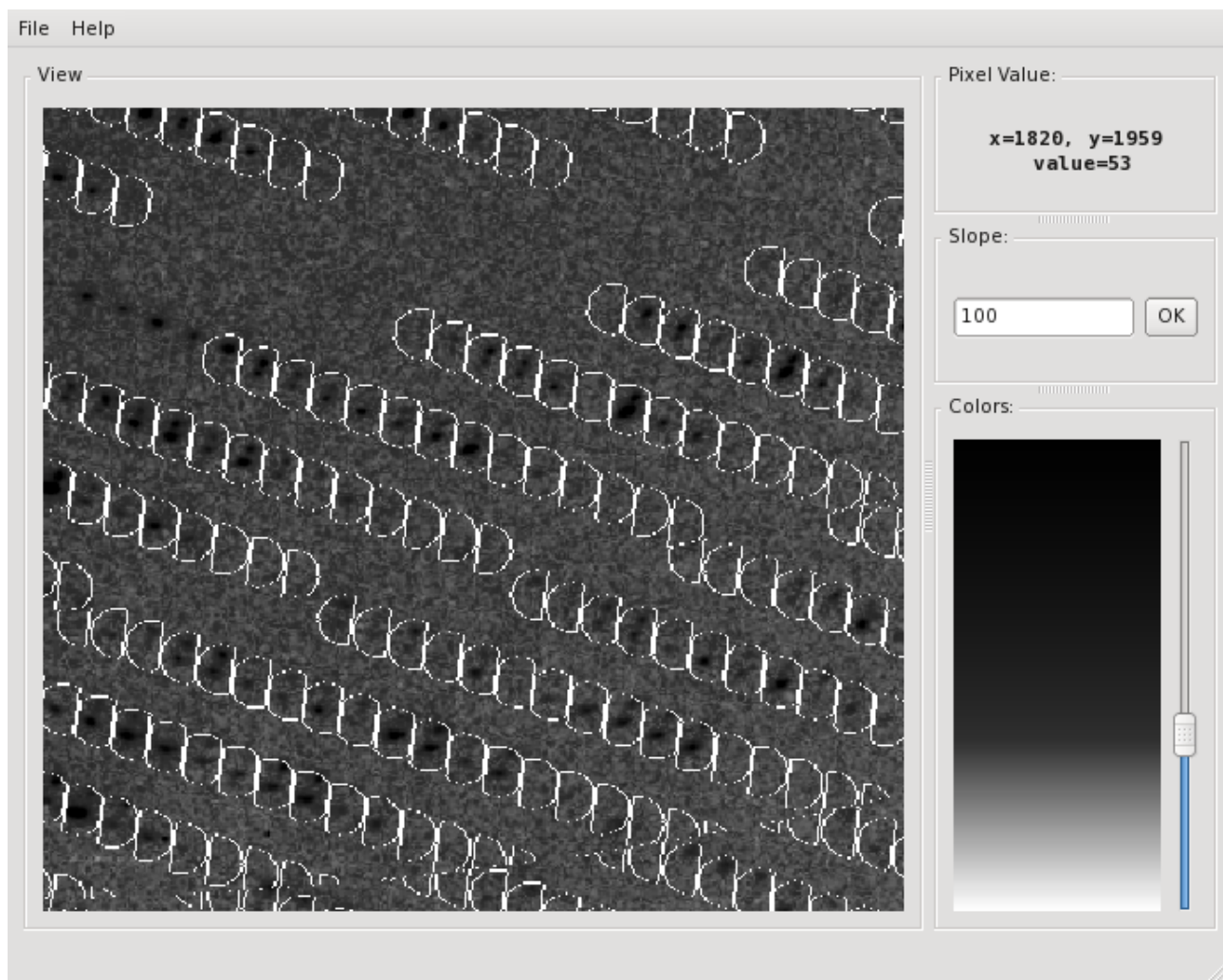
This is an excerpt from CORRECT.LP :

```

REFINED PARAMETERS:  DISTANCE BEAM ORIENTATION CELL AXIS
USING 166758 INDEXED SPOTS
STANDARD DEVIATION OF SPOT POSITION (PIXELS) 1.80
STANDARD DEVIATION OF SPINDLE POSITION (DEGREES) 0.61
CRYSTAL MOSAICITY (DEGREES) 0.557
DIRECT BEAM COORDINATES (REC. ANGSTROEM) 0.001590 -0.003616 1.021443
DETECTOR COORDINATES (PIXELS) OF DIRECT BEAM 1536.60 1519.03
DETECTOR ORIGIN (PIXELS) AT 1528.72 1536.95
CRYSTAL TO DETECTOR DISTANCE (mm) 370.85
LAB COORDINATES OF DETECTOR X-AXIS 1.000000 0.000000 0.000000
LAB COORDINATES OF DETECTOR Y-AXIS 0.000000 1.000000 0.000000
LAB COORDINATES OF ROTATION AXIS 0.999995 -0.001537 -0.002692
COORDINATES OF UNIT CELL A-AXIS -41.622 -128.332 80.636
COORDINATES OF UNIT CELL B-AXIS 45.889 72.744 139.459
COORDINATES OF UNIT CELL C-AXIS -551.193 220.474 66.370
REC. CELL PARAMETERS 0.006362 0.006103 0.001674 90.000 90.000 90.000
UNIT CELL PARAMETERS 157.174 163.848 597.351 90.000 90.000 90.000
E.S.D. OF CELL PARAMETERS 7.9E-01 8.7E-01 3.0E+00 0.0E+00 0.0E+00 0.0E+00
SPACE GROUP NUMBER 22
...
a b ISa
7.764E+00 7.144E-04 13.43
...
NOTE: Friedel pairs are treated as different reflections.
SUBSET OF INTENSITY DATA WITH SIGNAL/NOISE >= -3.0 AS FUNCTION OF RESOLUTION
RESOLUTION NUMBER OF REFLECTIONS COMPLETENESS R-FACTOR R-FACTOR COMPARED I/SIGMA R-meas Rmrgd
LIMIT OBSERVED UNIQUE POSSIBLE OF DATA observed expected
8.29 26331 7014 7090 98.9% 5.3% 5.5% 26314 20.22 6.3% 4.
5.90 48060 12542 12555 99.9% 8.4% 8.4% 48060 13.36 9.8% 8.
4.83 61534 16074 16144 99.6% 11.0% 10.6% 61534 10.69 12.9% 13.
4.19 71665 18994 19085 99.5% 12.5% 11.8% 71658 9.61 14.6% 16.
3.75 77668 21598 21677 99.6% 19.5% 19.3% 77668 6.33 23.0% 27.
3.42 78594 23767 23865 99.6% 28.3% 29.5% 78582 4.14 34.0% 45.
3.17 64135 24351 26036 93.5% 42.7% 46.4% 60830 2.18 52.9% 78.
2.97 40861 20207 27920 72.4% 63.8% 72.3% 35172 1.18 83.3% 118.
2.80 23238 15074 29715 50.7% 89.5% 104.9% 15359 0.63 122.8% 175.
total 492086 159621 184087 86.7% 14.8% 15.1% 475177 6.17 17.8% 30.
NUMBER OF REFLECTIONS IN SELECTED SUBSET OF IMAGES 501334
NUMBER OF REJECTED MISFITS 8429
NUMBER OF SYSTEMATIC ABSENT REFLECTIONS 0
NUMBER OF ACCEPTED OBSERVATIONS 492905
NUMBER OF UNIQUE ACCEPTED REFLECTIONS 159845

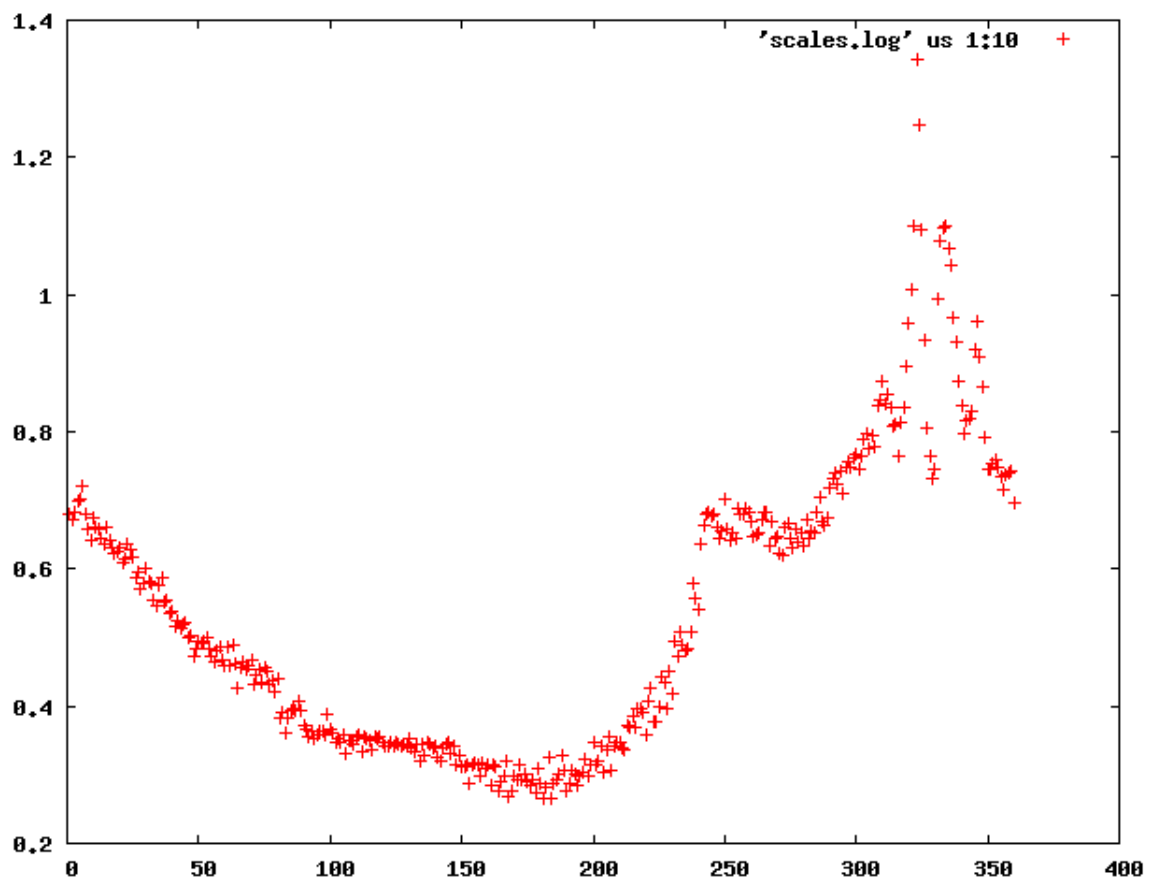
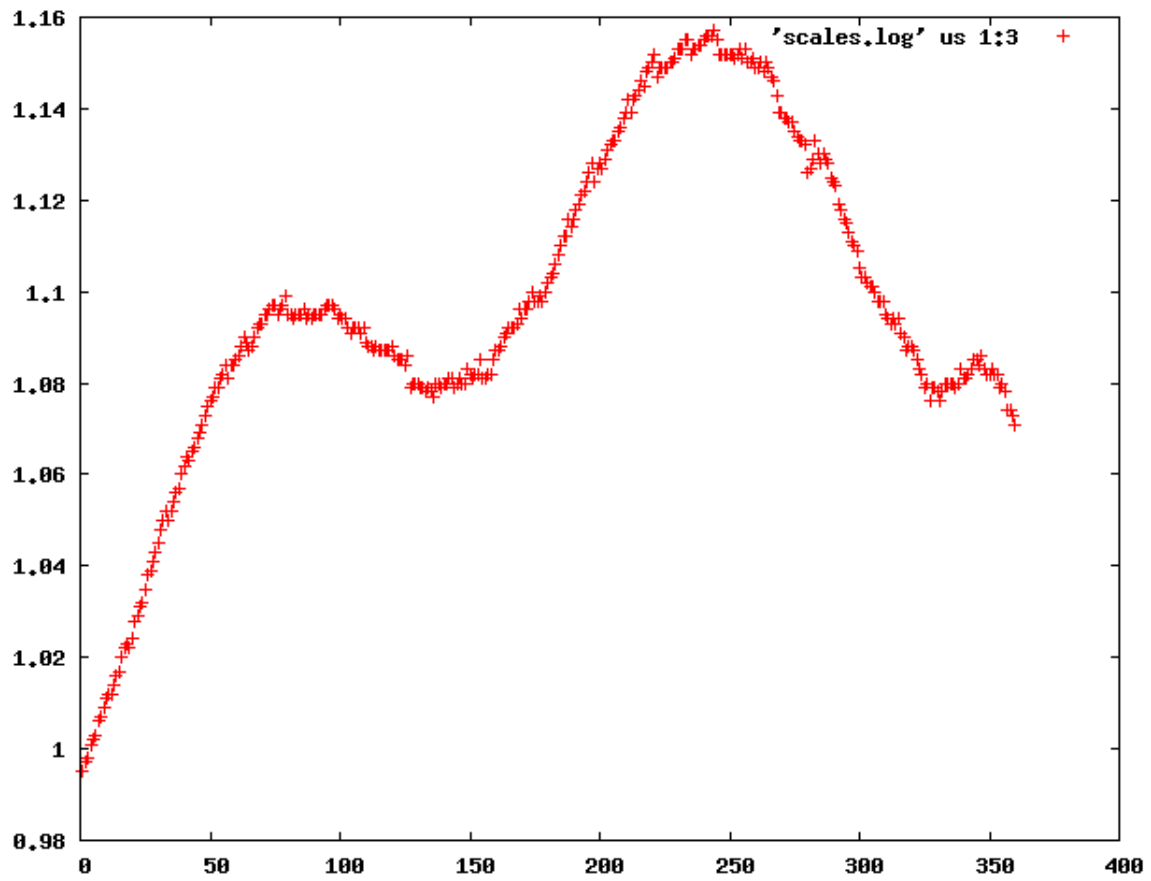
```

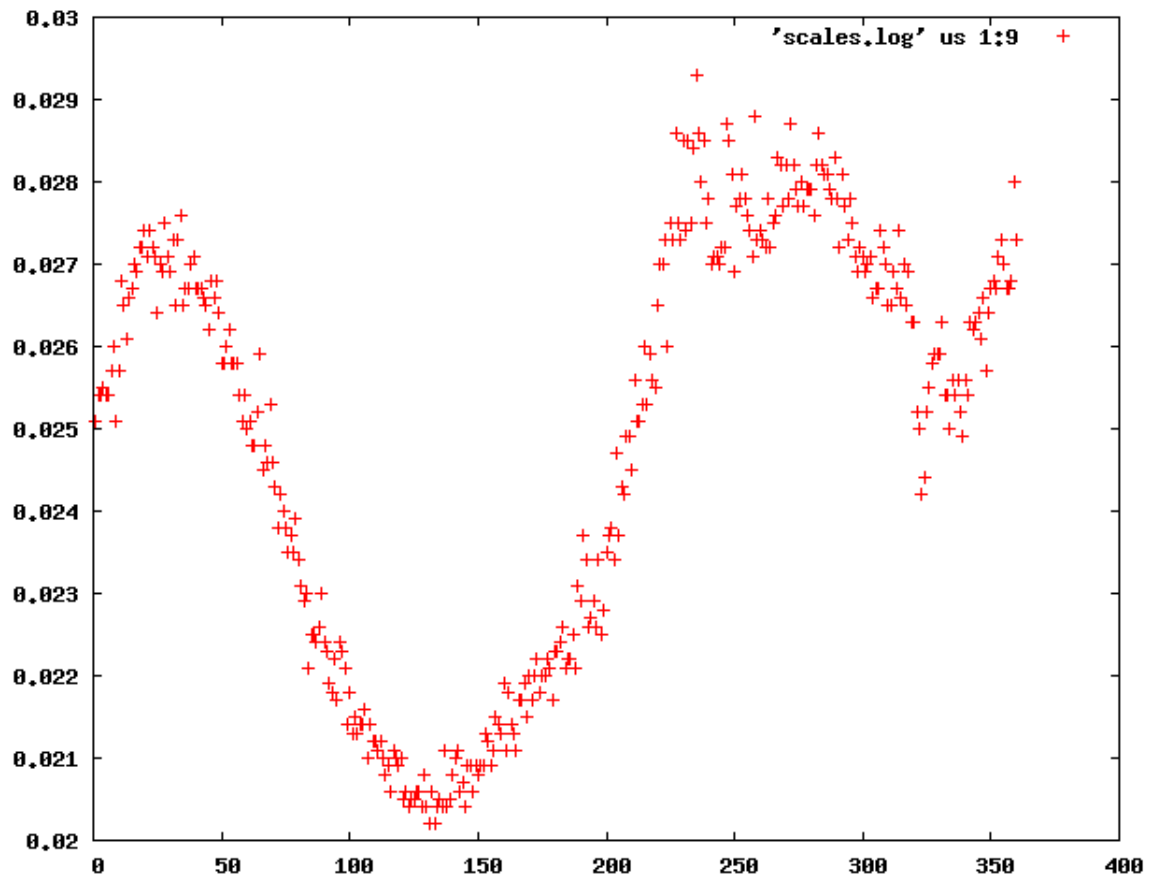
The verdict is clear: high mosaicity ($> 0.5^\circ$) bad ISa, anomalous correlation $> 30\%$ only to about 5 Å. The reason becomes clear if we load FRAME.cbf into XDS-Viewer, and zoom in:



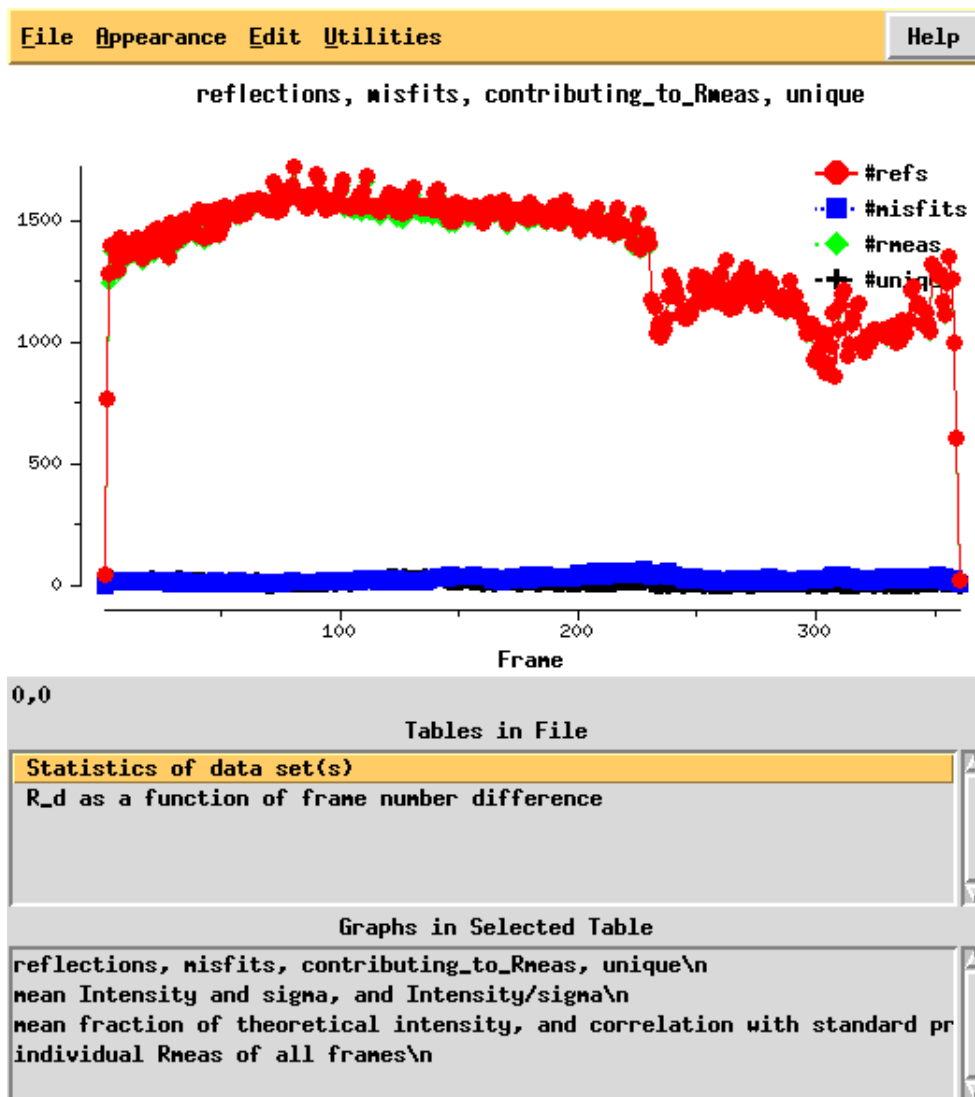
It is clear that these split reflections provide bad data. Fortunately it seems that in other areas of the detector, the reflections look better.

We can use the "scalefactors" jiffy to investigate the scale factor, and the estimates for mosaicity and beam divergence of each frame:

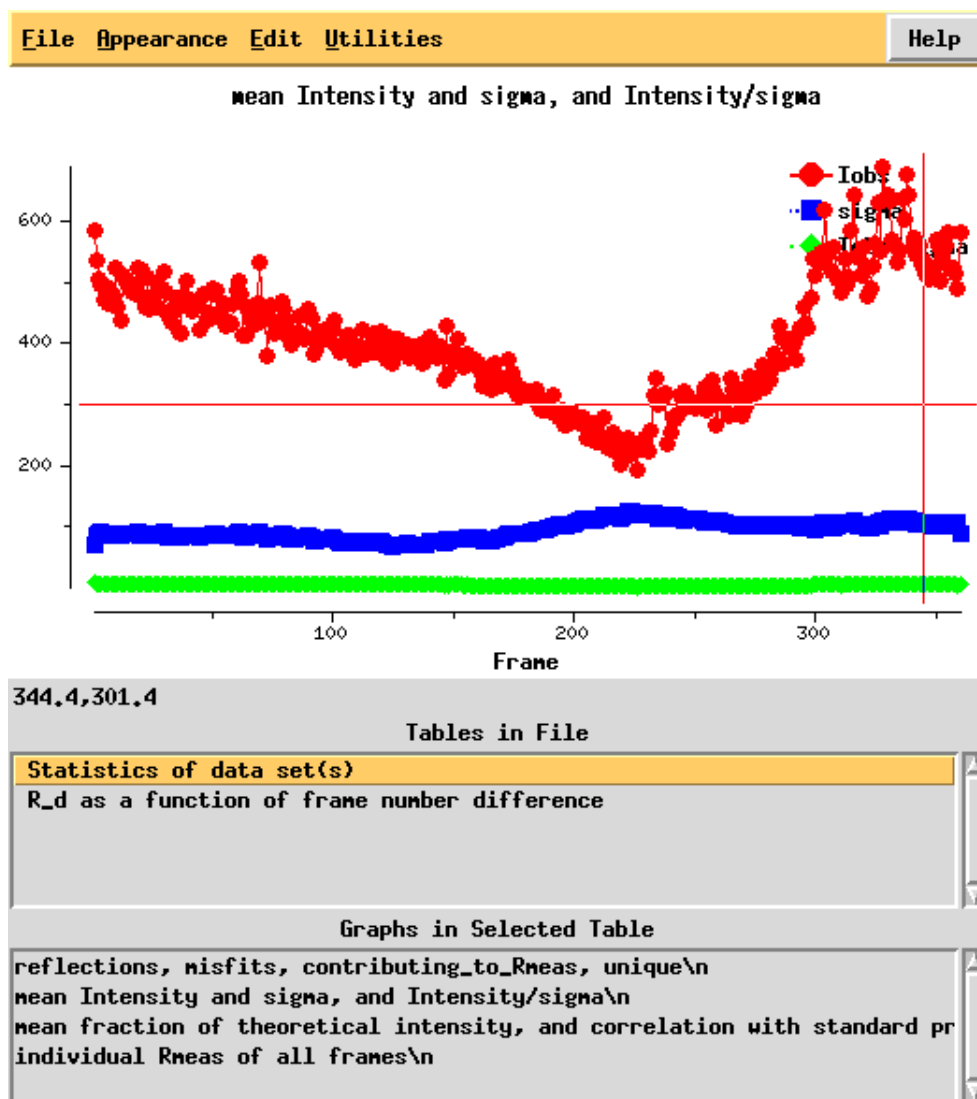




Next, we can use `xdsstat` to get frame-wise statistics from `XDS_ASCII.HKL`:



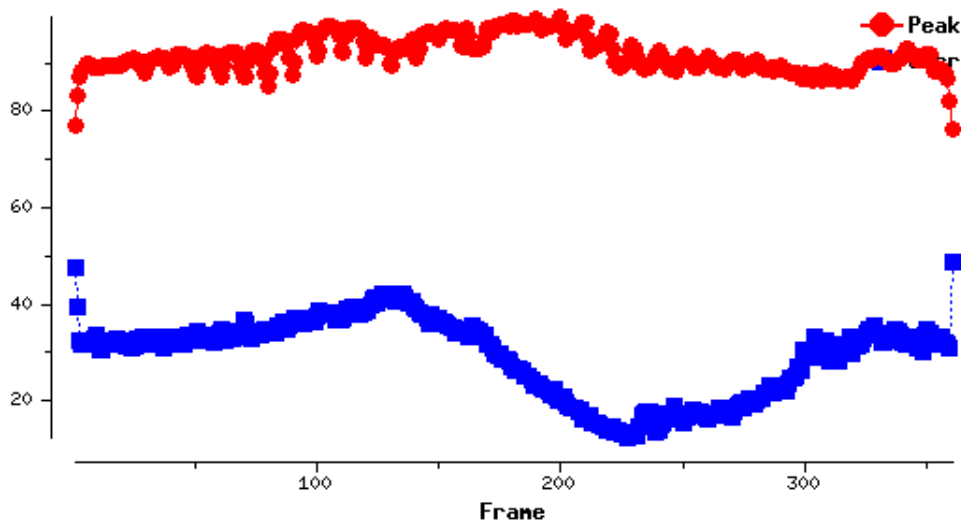
This shows a "jump" around frame 225 which is always bad for experimental phasing!



Around frame 225 the data are weakest, but they recover.

File Appearance Edit Utilities Help

mean fraction of theoretical intensity, and correlation with standard profi



278.0,5.939

Tables in File

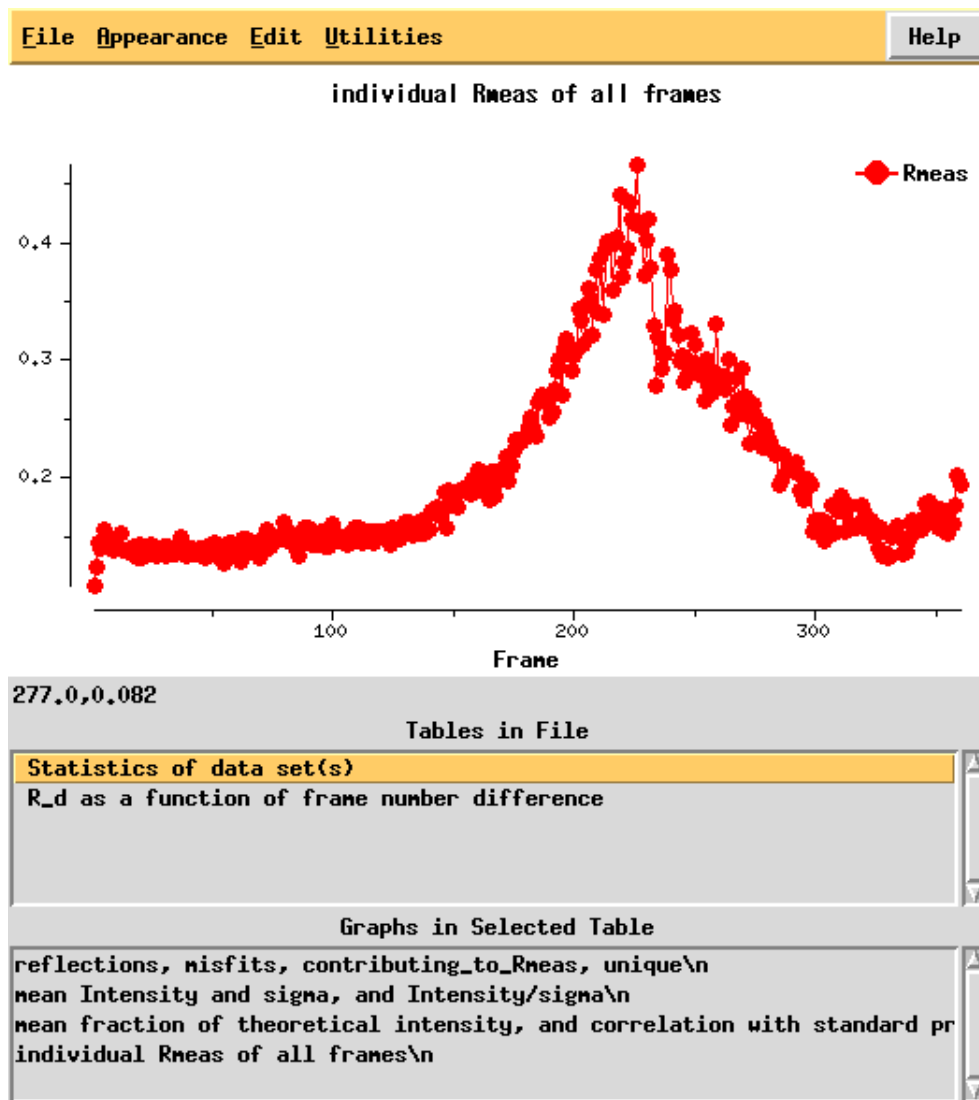
Statistics of data set(s)

R_d as a function of frame number difference

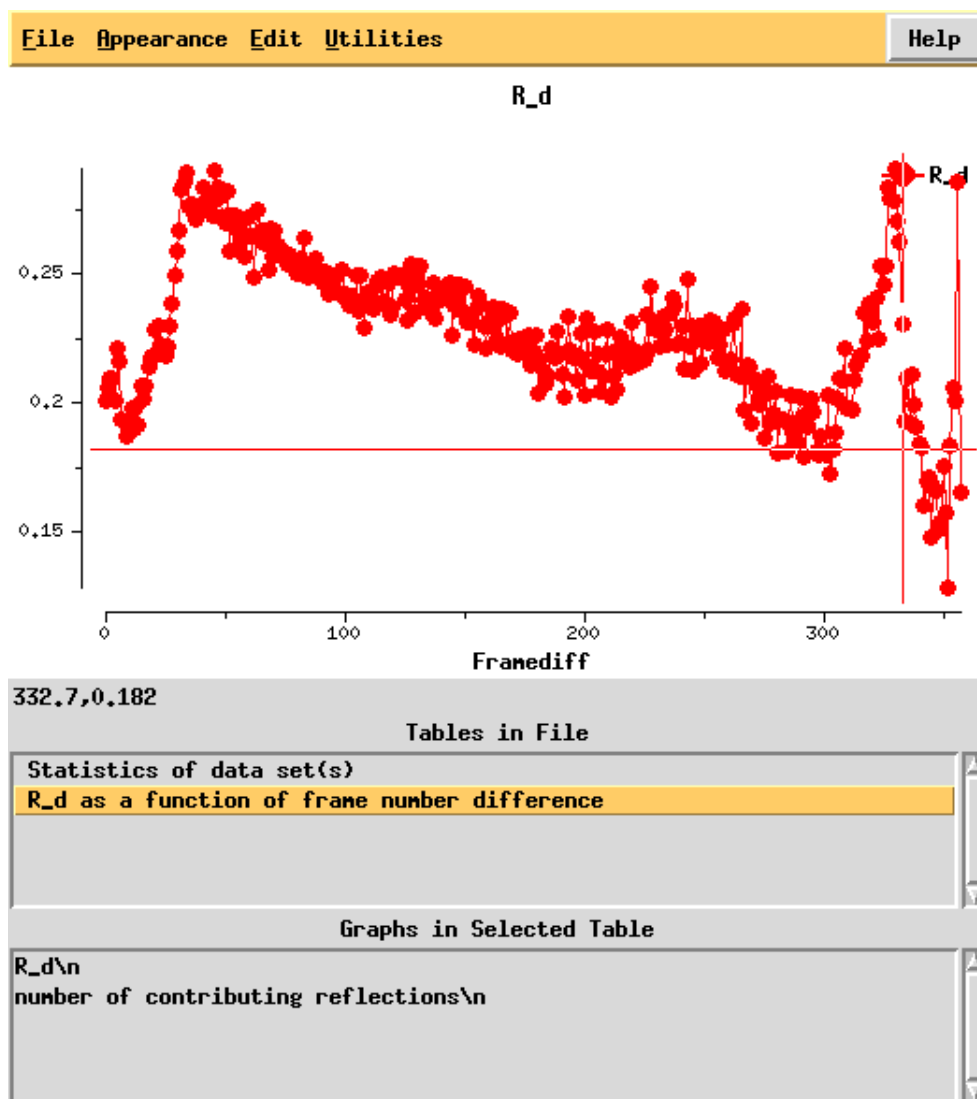
Graphs in Selected Table

reflections, misfits, contributing_to_Rneas, unique\
mean Intensity and sigma, and Intensity/sigma\
mean fraction of theoretical intensity, and correlation with standard pr
individual Rneas of all frames\

In particular the correlation against standard profiles (blue curve) is *really* low.



R-factors peak around frame 225.



R_d helps to quantify radiation damage. Unfortunately, for this dataset this "R-factor as a function of frame number difference" behaves wildly, so we cannot use 0-dose extrapolation, like we successfully did for 1Y13.

High-remote

Due to a beamline problem, high-remote data collection stopped after 269 frames of 0.5° (the final frame is already affected). After restart of the beamline, another 100 frames were collected but they later turned out to merge badly with the first 269 frames - a hint that the monochromator was still heating up, or similar. So the latter frames were left out. The 269 frames are here (<ftp://turn5.biologie.uni-konstanz.de/pub/datasets/3csl-hrem.tar>) (1.4 Gb).

From CORRECT.LP :

```

a      b      ISa
6.595E+00  3.032E-04  22.36
...

NOTE:      Friedel pairs are treated as different reflections.

SUBSET OF INTENSITY DATA WITH SIGNAL/NOISE >= -3.0 AS FUNCTION OF RESOLUTION
RESOLUTION      NUMBER OF REFLECTIONS      COMPLETENESS R-FACTOR      R-FACTOR COMPARED I/SIGMA      R-meas      Rmrgd
LIMIT           OBSERVED  UNIQUE  POSSIBLE  OF DATA  observed  expected
8.21            20187   6913   7245     95.4%    2.9%     3.3%    20065   27.89   3.6%    2.
5.84            36272   12417  12782    97.1%    5.5%     5.3%    36116   17.46   6.8%    7.
4.78            46716   16015  16500    97.1%    6.8%     6.5%    46473   14.43   8.4%    10.
4.15            55299   18949  19484    97.3%    7.5%     7.2%    55003   13.00   9.3%    11.
3.71            63751   21798  22065    98.8%    12.2%    12.2%   63371   8.51    15.1%   20.
3.39            70787   24180  24422    99.0%    19.5%    20.0%   70378   5.58    24.0%   33.
3.14            61197   25100  26452    94.9%    32.5%    34.2%   57925   2.88    41.3%   63.
2.94            40481   21869  28566    76.6%    53.5%    57.8%   33568   1.42    72.1%  112.
2.77            24584   16962  30228    56.1%    76.4%    82.4%   15055   0.77    107.5% 163.
total          419274  164203  187744   87.5%    10.3%    10.4%  397954   8.06    12.9%   24.

NUMBER OF REFLECTIONS IN SELECTED SUBSET OF IMAGES  428770
NUMBER OF REJECTED MISFITS                          9102
NUMBER OF SYSTEMATIC ABSENT REFLECTIONS             0
NUMBER OF ACCEPTED OBSERVATIONS                     419668
NUMBER OF UNIQUE ACCEPTED REFLECTIONS                164343

```

Inflection

360 frames (0.5° oscillation) at the inflection wavelength were collected after the peak data. They can be downloaded from here (<ftp://turn5.biologie.uni-konstanz.de/pub/datasets/3csl-ip.tar>) (1.8 Gb).

CORRECT.LP has:

```

a      b      ISa
6.514E+00  5.329E-04  16.97
...

NOTE:      Friedel pairs are treated as different reflections.

SUBSET OF INTENSITY DATA WITH SIGNAL/NOISE >= -3.0 AS FUNCTION OF RESOLUTION
RESOLUTION      NUMBER OF REFLECTIONS      COMPLETENESS R-FACTOR      R-FACTOR COMPARED I/SIGMA      R-meas      Rmrgd
LIMIT           OBSERVED  UNIQUE  POSSIBLE  OF DATA  observed  expected
8.28            26530   7039   7111     99.0%    4.4%     4.6%    26516   23.79   5.2%    3.
5.90            48700   12589  12598    99.9%    8.7%     8.5%    48700   14.00  10.2%    9.
4.83            62546   16177  16208    99.8%    12.0%    11.5%   62546   10.85  13.9%   14.
4.18            72644   19076  19117    99.8%    13.6%    13.1%   72637   9.64   15.9%   17.
3.75            80829   21710  21740    99.9%    23.8%    24.0%   80829   5.92   27.9%   32.
3.42            86652   23874  23917    99.8%    38.5%    39.9%   86652   3.71   45.2%   53.
3.17            73630   25264  26115    96.7%    64.4%    68.0%   71837   1.82   78.5%  109.
2.96            48079   22582  28004    80.6%    99.2%   107.3%  43325   0.86  129.6%  186.
2.80            28417   17801  29828    59.7%   155.6%  168.5%  20099   0.43  214.2%  307.
total          528027  166112  184638   90.0%    17.6%    17.8%  513141   5.98   20.9%   38.

NUMBER OF REFLECTIONS IN SELECTED SUBSET OF IMAGES  534898
NUMBER OF REJECTED MISFITS                          6486
NUMBER OF SYSTEMATIC ABSENT REFLECTIONS             0
NUMBER OF ACCEPTED OBSERVATIONS                     528412
NUMBER OF UNIQUE ACCEPTED REFLECTIONS                166224

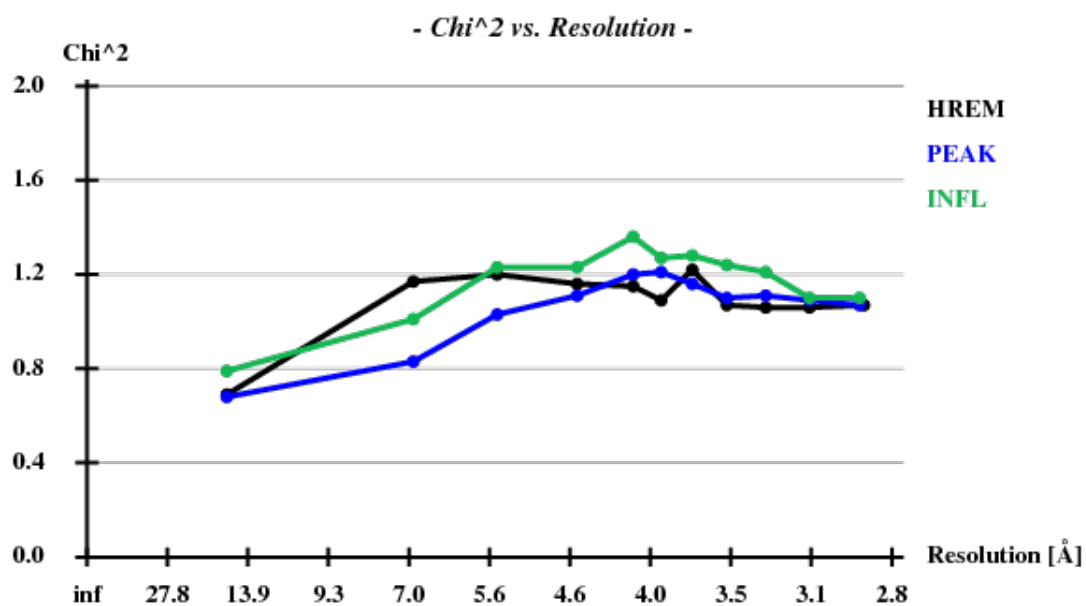
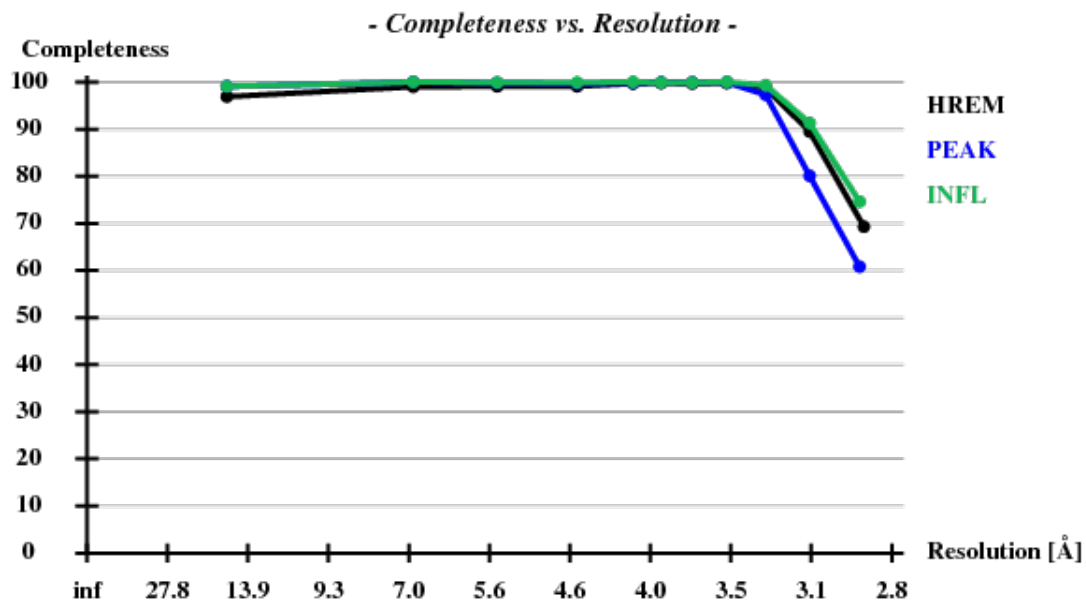
```

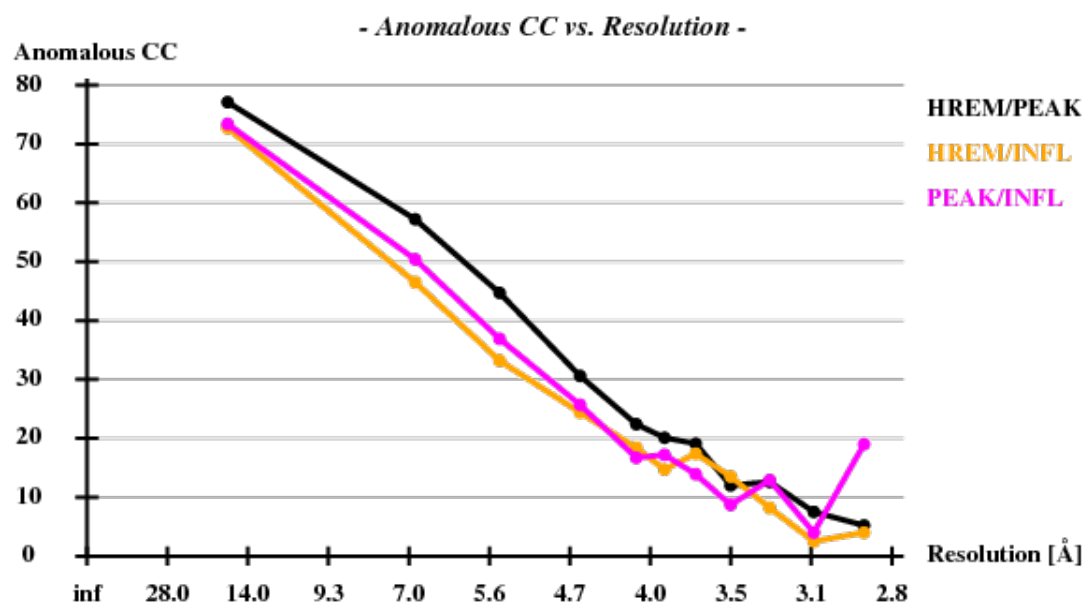
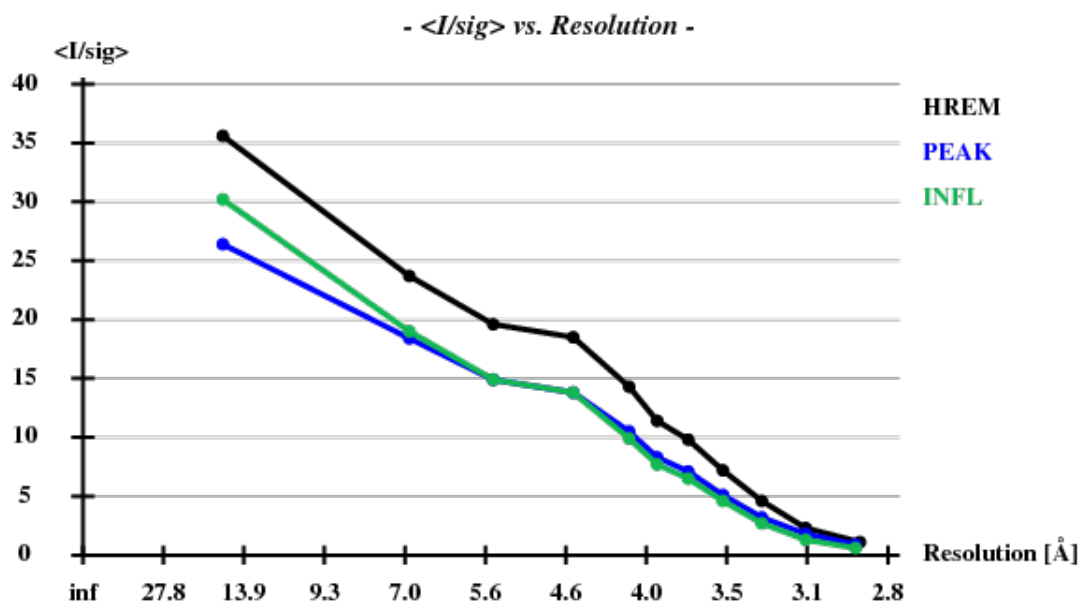
Scaling

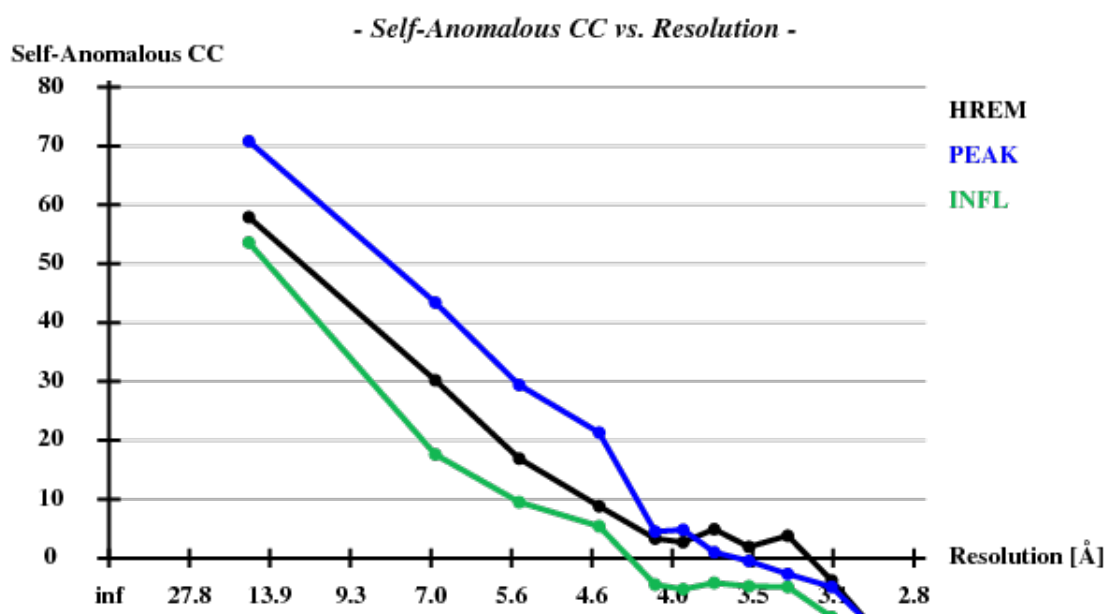
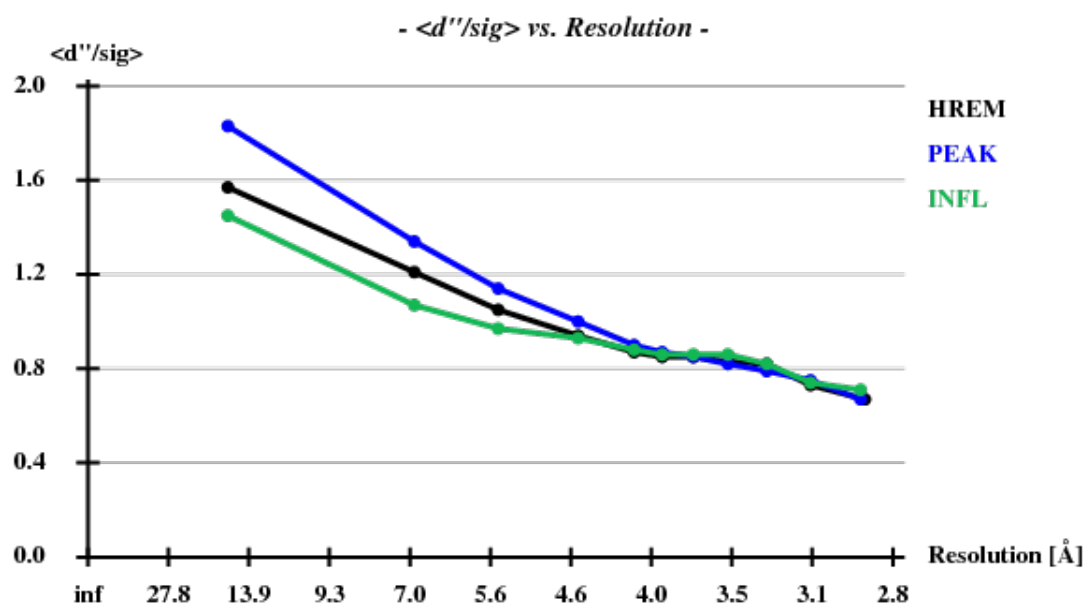
Structure solution

We use hkl2map for solving the structure.

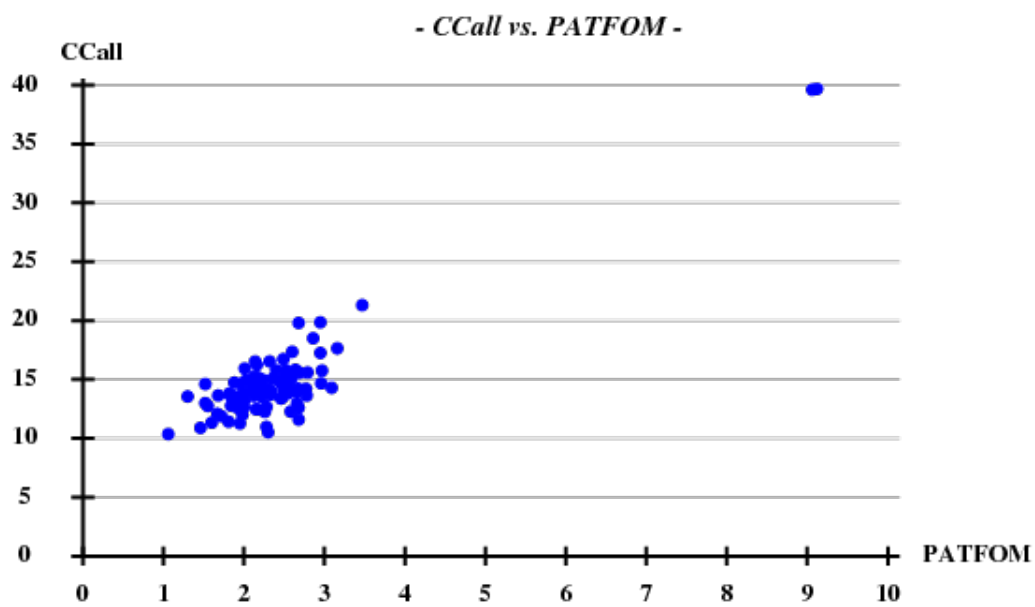
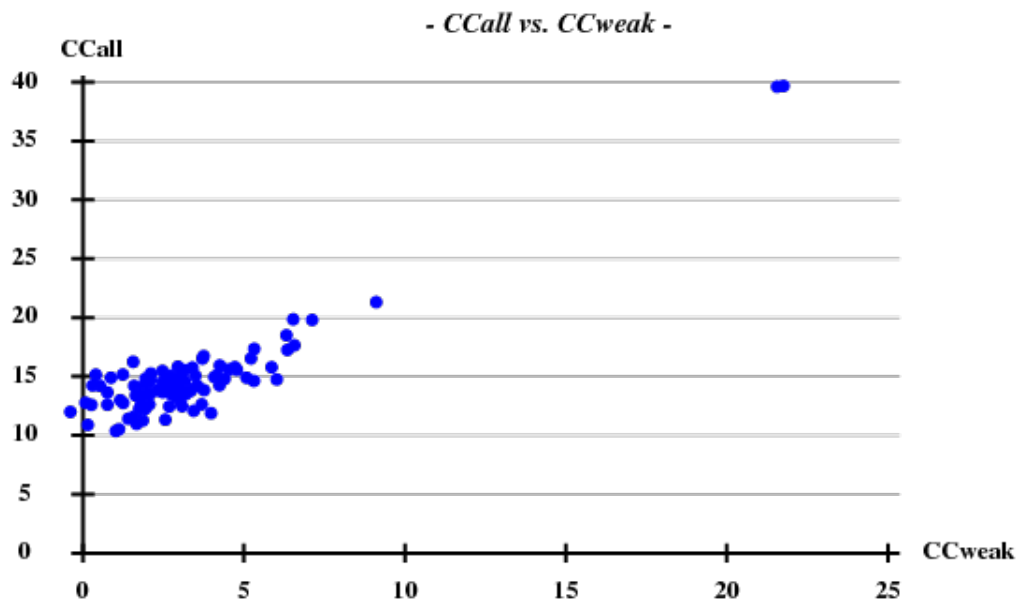
SHELXC statistics

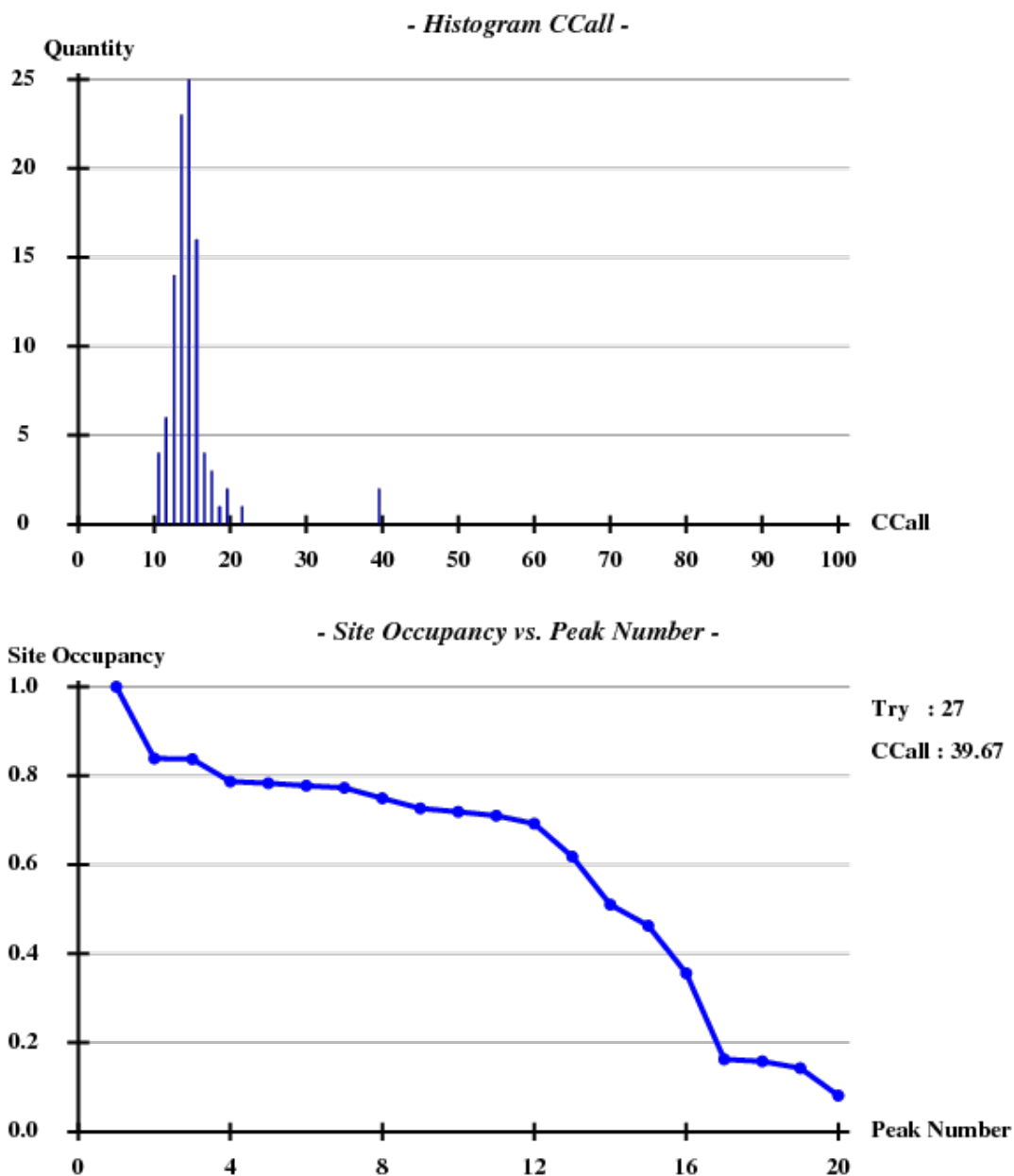






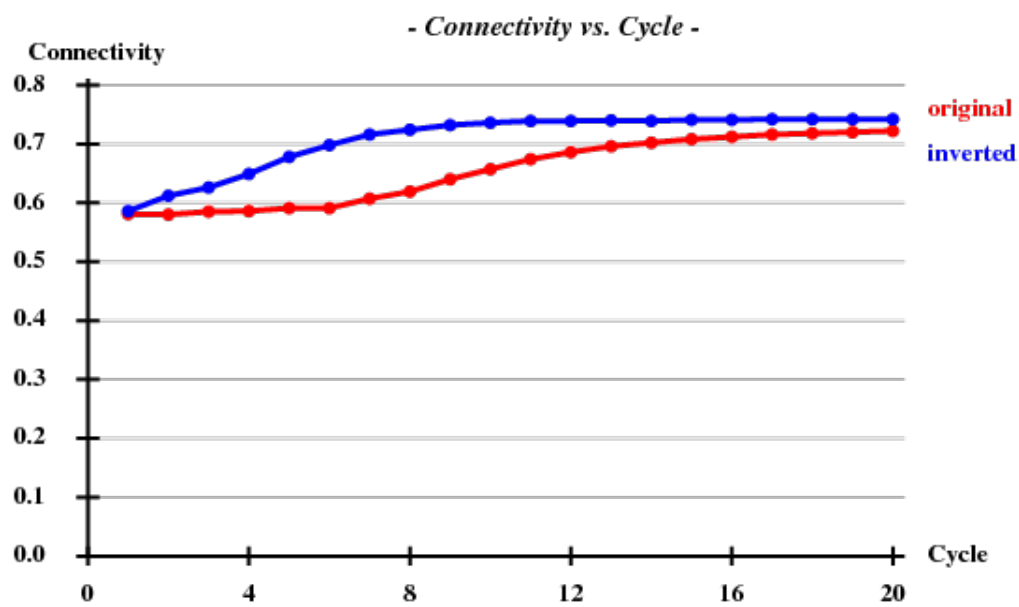
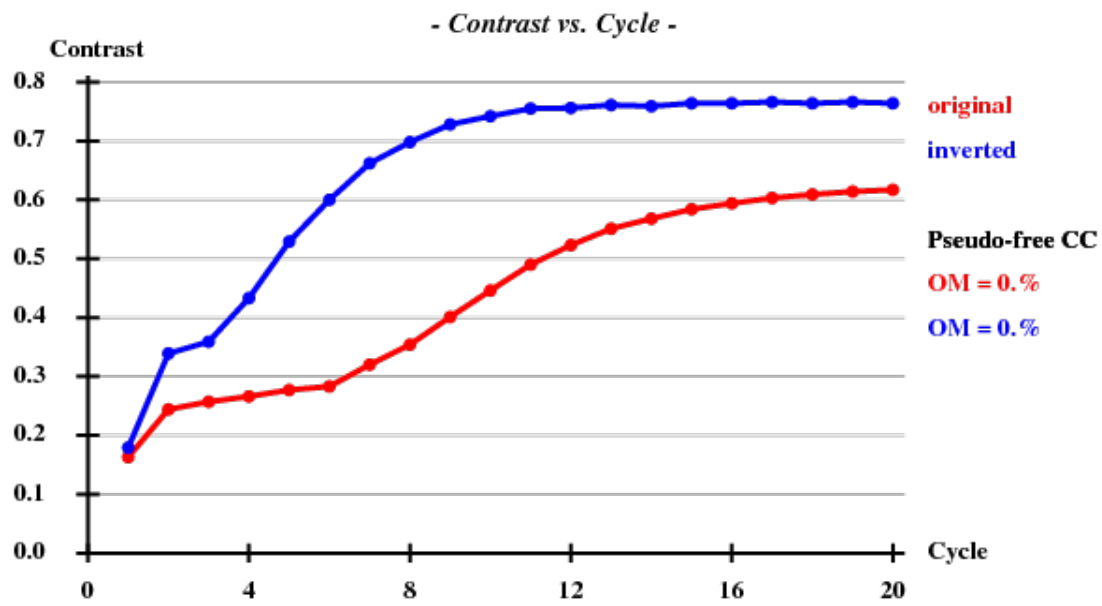
SHELXD statistics

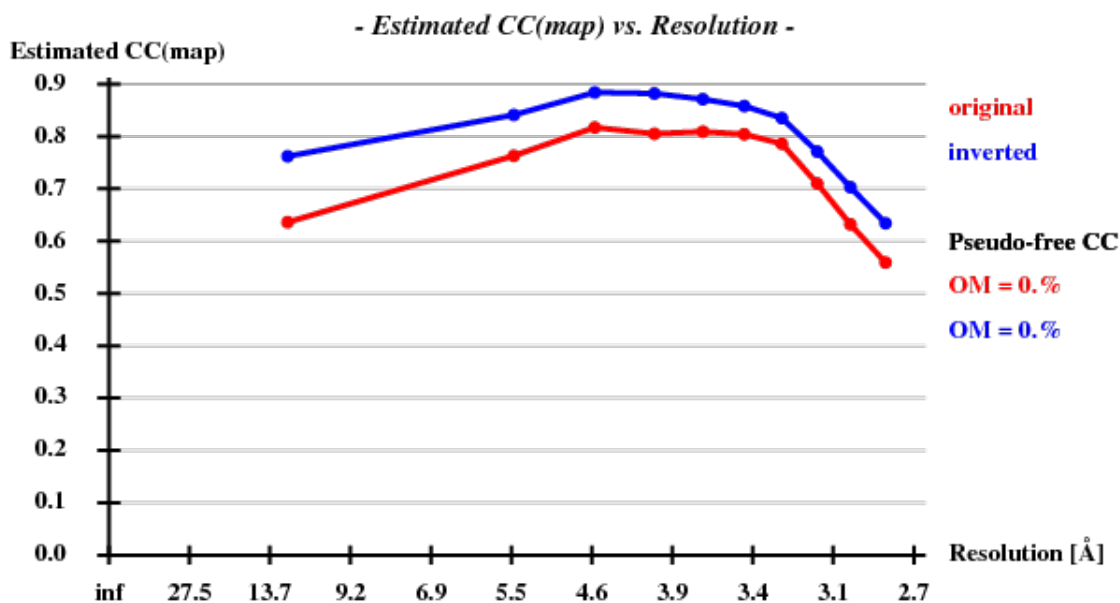




SHELXE statistics

Since there are 1852 residues in the ASU, the solvent content is about 72.5%. The correct hand ("inverted") becomes immediately clear - it is superior in all respects than the "original" hand.

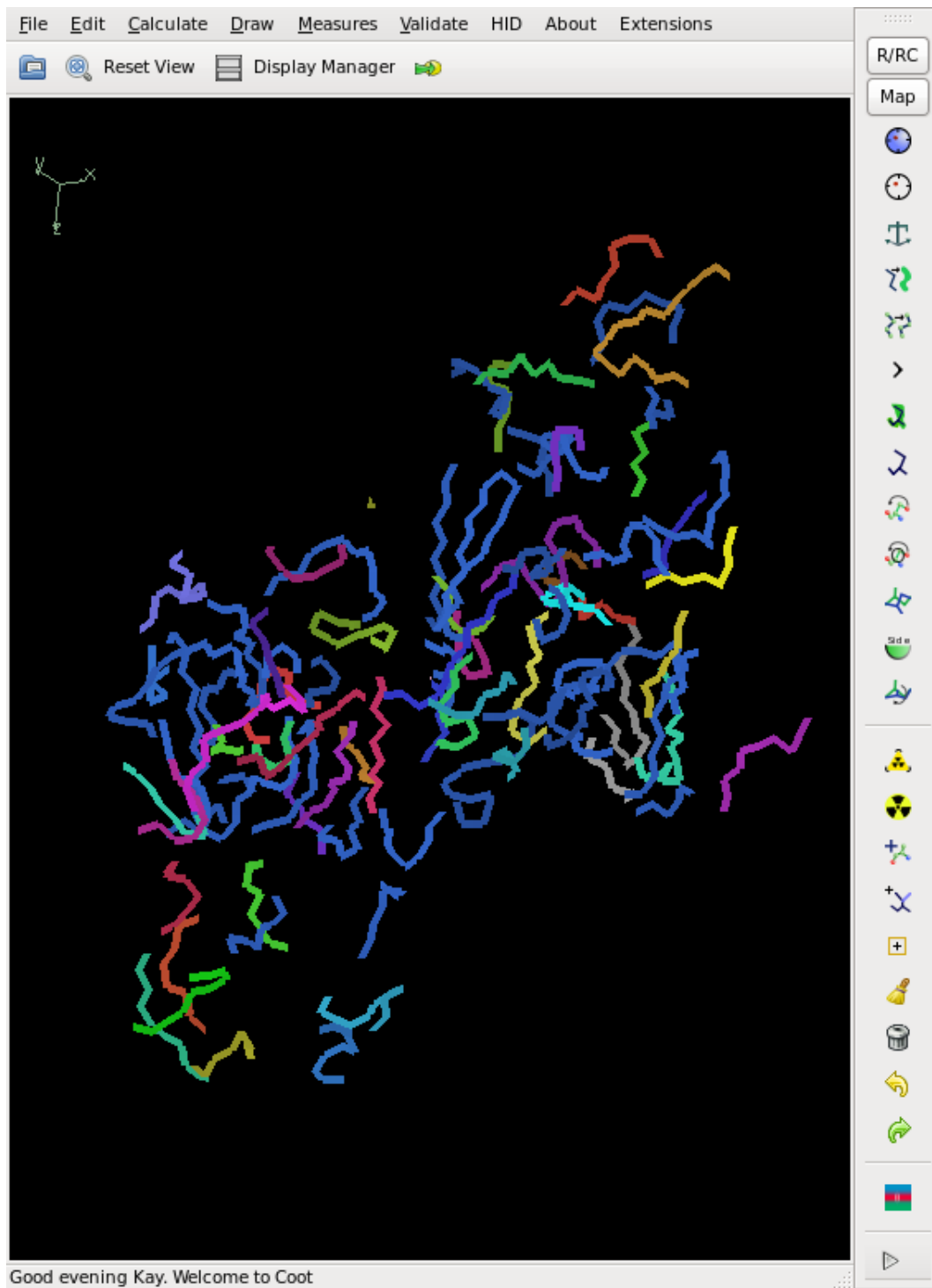


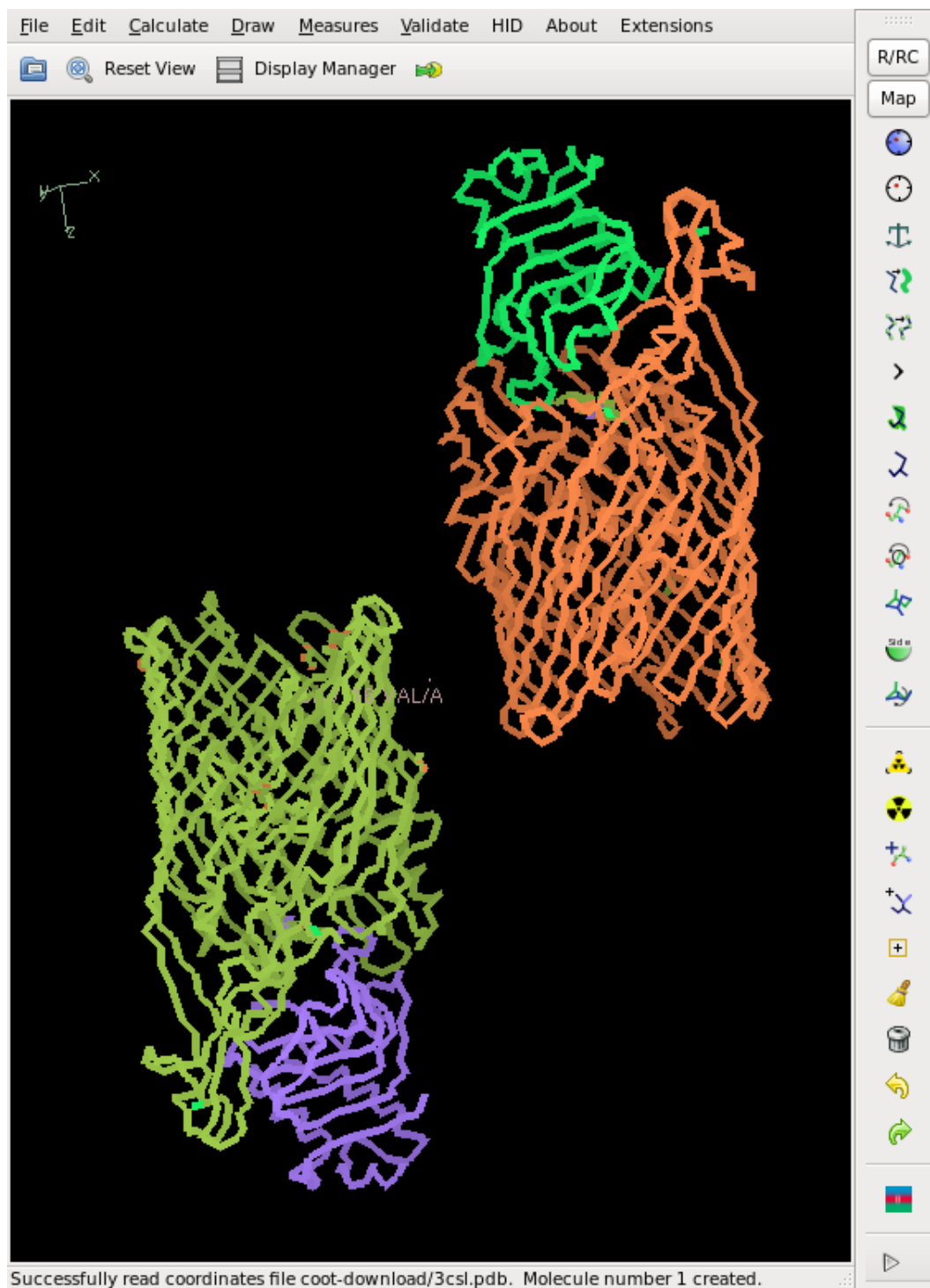


Then we try to get a poly-ala backbone tracing, using the 16 Se sites found by SHELXD:

```
shelxe.beta -a -q -m15 -s0.725 -b -h16 -n2 -i mad mad_fa
```

This does not yield a complete chain, but rather about 50% of it, and the CC is slightly less than 20%, so that we cannot consider the structure as solved yet. However, the phases are good enough for finding 2 additional sites. We iterate this, and finally the Se sites all have a density of at least 20 sigma. The model from shelxe as well as the deposited structure are shown below:





No doubt that one can solve the structure from here, maybe after HA refinement with SHARP, and model building with buccanner or Arp/wArp.

Availability of data

There are files with amplitudes (3csl-pk-F.mtz, 3csl-rh-F.mtz, 3csl-ip-F.mtz) and intensities (3csl-pk-I.mtz, 3csl-rh-I.mtz, 3csl-ip-I.mtz) as well as mad_i.pdb and mad_i.phs (written by SHELXE) available from [1] (<ftp://turn5.biologie.uni-konstanz.de/pub/xds-dated/3csl/>) .

additional information for those who want to complete the structure

These are the entire sequences of HasR and HasA - before solving the structure it was not known that the N-terminus of HasR was disordered.

```

AQAEASSAQAQQKNFNIAAQPLQSAMLRFAEQAGMQVFFDEVKLDGMQAAALNGSMSVEQGLRRLIGGNPVAFRLLQPQGGQIVLSRLPTANGDGGALALD
SLTVLGAGGNANDWVYDEPRSVSVISREQMDNRPARHAADILEQTTGAYSSVSQQDPALSVNIRGIQDYGRVNMNIDGMRQNFQKSGHGQRNGTMYIDS
ELLSGVTIDKGTGGMGSAAGTLGGIATFNTVSASDFLAPGKELGGKLHASTGDNGTHFIGSGILALGNETGDILLAASERHLGDYWPGNKGDIGNIRINN
DTGNYDRYAESIKNNKIPDTHYRMSRLAKVGWNL PANQRLQLSYLQTQTASPIAGTLNLGTRPPYELGWKRTGYTDMARNAAFDYS LAPEDVDWLD
QAKLYYVDTQDDSDTYSTSSLLDNGYATRRLRRTYGAQAQNTSRFSLAPGHDFRANYGLEFYDKATSDSSRQMEGVT PAGNRSVASL FANLTYDYDGW
LTLEGLRYDRYRLRGQTGLSYDPLAKDGQRYTIDNPCKALRLTGCSTTTREDWDVDRDQGKLSPTLAVAVRPGVLEWLELYTTYGKSWRPPAITETLTNG
SAHSSSTQYPNPFLLQERSRAWVGFVQPDWLFEGDRLVAKVaYFDTKVDNYINLAIDRNKPGLVQPSIGNAAYVNNLSKTRFRGLEQLNYDAGVFY
ADLTYTHMIGKNEFCSNKAWLGGRLRYGDGSRGFYVEPDAASNDVFTCDGGTQFGSAAYLPGDRGSVTLGGRAFDRKLDAGVTVRFAPGYQDSSVPSN
YPYLADWPKYTLFDLYASYKLTDSLTLRGSVENLTNRAYVVSYGETLANTLGRGRTVQGGVEYRF
MRGSHHHHHHGIRMRARYPAFSVNYDSSFYGGYSIHLYLQGWASTFGDVNH
TNGNVTDANSGGFYGGSLSGSQYAISSSTANQVTAFAVAGGNLTYTLFNEPA
HTLYGQLDSL SFGDGLSGGDTSPYSIQVPDVSFGGLNLSSLQAQGHGQV
HQVYGLMSGDTGALETALNGILDDYGLSVNSTFDQVAAATAVGVQHADS
PELLAA

```

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- This page was last modified on 18 March 2011, at 21:08.